

Stochastically forced long-range interacting systems

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Statistical Mechanics of self-gravitating particles

Fondation des Treilles, 25 October 2012

joint work with **F. Bouchet**, **S. Gupta**, **T. Dauxois** and **S. Ruffo**

CN, S. Gupta, S. Ruffo, T. Dauxois and F. Bouchet, Letter to JSTAT, L01002 (2012)

CN, S. Gupta, S. Ruffo, T. Dauxois and F. Bouchet, arXiv:1210.0492, submitted to JSTAT

d : spatial dimensions of the systems

$$r \gg r_{\text{typ}} \quad v(r) \sim \frac{1}{r^\alpha} \quad \alpha \leq d$$

- Gravitational system, one-component plasma

$$H = \sum_{i=1}^N \frac{p_i^2}{2m} + \frac{1}{2N} \sum_{i,j=1}^N v(q_i - q_j)$$

- 2D, Quasi-2D turbulence and geophysical flows

- ex.: 2-d Euler:

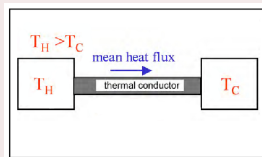
$$\partial_t \omega + \mathbf{v} \cdot \nabla \omega = 0 \quad \text{with} \quad \omega = \nabla \wedge \mathbf{v}, \quad \omega = \Delta \psi$$

$$\mathcal{E} = - \int d\mathbf{r} d\mathbf{r}' \omega(\mathbf{r}) G(\mathbf{r}, \mathbf{r}') \omega(\mathbf{r}'), \quad G \sim \log |\mathbf{r} - \mathbf{r}'|$$

- Quasi-2D models: Shallow-water equation, Quasi-geostrophic equation, ...

- ...

Short-range systems



Non-equilibrium stationary states sustained by flux of conserved quantities,
broken detailed balance, ...
no analogous to Boltzmann-Gibbs theory

Long-range systems

forcing can act coherently on all the degrees of freedom:

- imposed electric fields on a plasma ?
- gravitational fields created by other galaxies ?
- ... ?

Fluid models

- 2D and Quasi-2D turbulence models
- Large scale structures in contrast with 3D turbulence

Geophysical systems

- Energy injection
 - Wind on oceans
 - Different layers in atmosphere
 - ...
- Dissipation on large scales
 - effect of boundaries



(Weak?) FLUX OF ENERGY from SMALL to LARGE SCALES
OUT OF EQUILIBRIUM PHENOMENA!

Stochastically forced quasi-Geostrophic equations

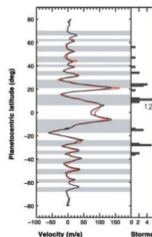
$$\partial_t q + \mathbf{v} \cdot \nabla q = -\alpha q + F(\mathbf{r}, t)$$

$$\langle F(\mathbf{r}, t) F(\mathbf{r}', t') \rangle = C(\mathbf{r} - \mathbf{r}') \delta(t - t')$$

q: quasi-geostrophic potential vorticity



Jupiter atmosphere



Jupiter Zonal wind (Voyager and Cassini, from Porco et al 2003)

IN THIS TALK: stochastically forced Long-range PARTICLE systems

Work in progress (with F. Bouchet & T. Tangarife)

similar theoretical techniques for stochastically forced 2d fluids

$$\partial_t \omega + \mathbf{v} \cdot \nabla \omega = -\alpha \omega + F(\mathbf{r}, t)$$

Similarity between 2d Euler and Vlasov equation

- non-linear transport equations
- infinite number of conserved quantities (Casimirs)
- Hamiltonian structure

- 1 Isolated long-range particle systems
- 2 Stochastically forced particle systems: the model
- 3 Kinetic theory
- 4 Comparison between kinetic theory and numerical simulations
- 5 Bistability

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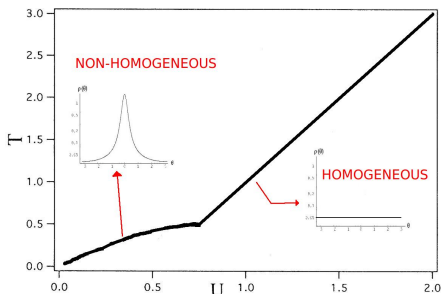
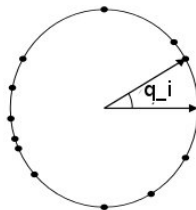
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Hamiltonian Mean Field model

$$\begin{aligned}
 H &= \sum_{i=1}^N \frac{p_i^2}{2} - \frac{1}{2N} \sum_{i,j} \cos(q_i - q_j) = \\
 &= \sum_{i=1}^N \frac{p_i^2}{2} - \frac{N}{2} |m|^2
 \end{aligned}$$



magnetization

$$m(t) = \sqrt{m_x^2 + m_y^2}$$

$$m_x(t) = \frac{1}{N} \sum_{i=1}^N \cos q_i$$

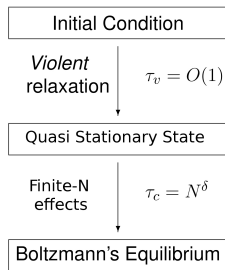
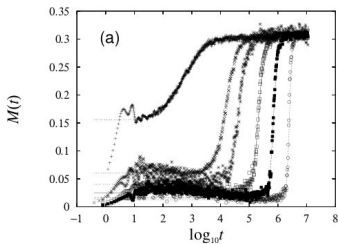
$$m_y(t) = \frac{1}{N} \sum_{i=1}^N \sin q_i$$

Computational cost of Molecular Dynamics $\sim N$

Isolated long-range systems: relaxation to equilibrium

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Relaxation to equilibrium

- Quasi-stationary states
 - Non-ergodicity
 - lifetime $\sim N$ (diverging with the system size!)

small parameter: $1/N$

$f(q, p, t)$: density in (q, p) at time t

Vlasov equation ($t \ll \tau_c \sim N^\delta$, $\delta > 0$)

$$\frac{\partial f}{\partial t} + p \frac{\partial f}{\partial q} - \frac{\partial f}{\partial p} \frac{\partial \Phi}{\partial q} = 0 \quad \Phi(q) = \int dq' v(q - q') f(q')$$

- Mean field approximation
 - exact for $N \rightarrow \infty$
- Quasi-Stationary States: stable equilibria

Infinite number of QSS

Lenard-Balescu equation ... ($t \sim \tau_c$)

$$\frac{\partial f}{\partial t} + p \frac{\partial f}{\partial q} - \frac{\partial f}{\partial p} \frac{\partial \Phi}{\partial q} = \frac{1}{N} C[f] \quad \text{Lowest order description of finite-N effects}$$

- weak correlations cause slow evolution
 - analogous to Boltzmann equation

slow relaxation through Quasi-Stationary-States

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STOCHASTIC EQUATIONS OF MOTION

$$\dot{q}_i = \frac{\partial H}{\partial p_i} \quad \dot{p}_i = -\frac{\partial H}{\partial q_i} - \alpha p_i + \sqrt{\alpha} F(q_i, t)$$

- F Gaussian stochastic process with $\langle F(q, t) \rangle = 0$

$$\langle dF(q, t) dF(q', t') \rangle = B(|q - q'|) \delta(t - t') dt$$

α : forcing and dissipation parameter

Coherent stochastic forces

- Coherent stochastic term: NOT $F_i(q_i, t)$
 - external stochastic field

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"Fourier expansion" of $F(q, t)$

$$g_k = \frac{1}{L} \int dq B(q) e^{-ikq} > 0, \quad F(q, t) = \sum_k g_k e^{ikq} W_k(t)$$

W_k : independent Wiener processes

g_k : forcing at the spatial scale $1/k$

Kinetic energy in a steady state: KINETIC TEMPERATURE

$$T = 2 \langle K_{ss} \rangle = \frac{1}{2} \sum_k g_k^2$$

Detailed balance $\leftrightarrow g_k = g \quad \forall k$

Can be far from detailed balance also for $\alpha \ll 1$

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- $N \gg 1$: number of degrees of freedoms
 - Plasma, self-gravitating systems
- $\alpha \ll 1$: weak forcing limit
 - Technical reason: small parameter

Timescales

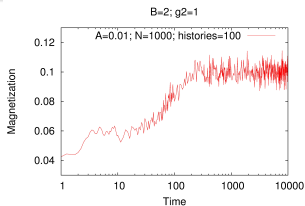
- Collective timescale: $\tau_c \sim N$
- Stochasticity: $\tau_s = 1/\alpha$

Limits

- Continuous limit: $N\alpha \gg 1$
 - Negligible finite size effects:
similar to 2D fluid models!
- $N\alpha \sim 1$ or $\ll 1$: simple generalization

Fluctuations of intensive observables

- Finite-size effects: $\sim 1/\sqrt{N}$
- Stochasticity: $\sqrt{\alpha}$



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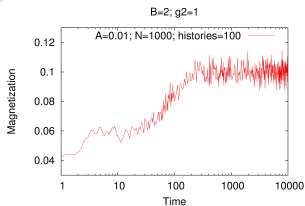
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Perturbation theory in the SMALL PARAMETER: $\alpha \gg 1/N$ FOKKER-PLANCK for the N -particles distribution function f_N

$$\frac{\partial f_N}{\partial t} + \text{Liouville terms} = - \sum_{i=1}^N \frac{\partial(\alpha p_i f_N)}{\partial p_i} - \frac{\alpha}{2} \sum_{i,j=1}^N C(q_i - q_j) \frac{\partial^2 f_N}{\partial p_i \partial p_j}$$

$$\langle F(q, t) F(q', t') \rangle = C(|q - q'|) \delta(t - t')$$

Exact: too much information for a macroscopic description

Distribution functions

- $f_s = \int f_N d[s+1] \dots d[N]$
- $f_1 = f$: density in μ -space
- f_2 : 2-particles correlations
-

 \Rightarrow

BBGKY hierarchy

$$\partial_t f_s = L[f_s, f_{s+1}]$$

How to close the BBGKY hierarchy?

First equation of the BBGKY

$$\frac{\partial f}{\partial t} + \text{Vlasov} - \frac{\partial}{\partial p_1} [\alpha p_1 f] - \alpha T \frac{\partial^2 f}{\partial p_1^2} = C[f_2]$$

- blue: order 1
- red: order α

non negligible two-particle correlations

Analogy with finite size effects in Hamiltonian systems

CORRELATIONS	finite- N	stochastic forces
	<ul style="list-style-type: none"> • Minimal project: discard 3-particles correlations 	
METHOD	<ul style="list-style-type: none"> • Solve the II eq. of BBGKY • Plug the result in the I eq. of BBGKY 	
KIN. EQ.	Lenard-Balescu	"Our" kin. eq.

Hypothesis

f stationary stable solution of Vlasov equation at every time

- -- > Time-scale separation $\sim f$ evolves slowly w.r.t. g

homogeneous system: $f(p, q, t) = f(p, t)$

- -- > explicit form of the kinetic equation

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connected components of correlations

$$f_2(1, 2) = f(1)f(2) + \alpha g(1, 2)$$

$$f_3(1, 2, 3) = f(1)f(2)f(3) + f(1)g(2, 3) + \dots + \dots + h(1, 2, 3)$$

...

it is SELF-CONSISTENT to suppose:

- $f \sim O(1)$
- $\alpha g \sim O(\alpha)$
- $h \sim \ll \alpha g$
- ...

Lowest order possible scheme if we want to describe the effect of the forcing

Discard three-particle and higher order correlations while taking into account two-particle correlations

- Analogous to derivation of Boltzmann eq. or Lenard-Balescu eq.

I equation BBGKY

$$\frac{1}{\alpha} \frac{\partial f}{\partial t} + \frac{1}{\alpha} \text{Vlasov} - \frac{\partial}{\partial p_1} [p_1 f] - T \frac{\partial^2 f}{\partial p_1^2} = \frac{\partial}{\partial p_1} \int d[2] v'(q_1 - q_2) g(1, 2, t)$$

II equation BBGKY

$$\frac{\partial g}{\partial t} + L_f^{(1)} g + L_f^{(2)} g = C(|q_1 - q_2|) \frac{\partial}{\partial p_1} \frac{\partial}{\partial p_2} f(1, t) f(2, t)$$

$L_f g$: linearized Vlasov operator around f acting on h

Time-scale separation

If f is a STATIONARY AND STABLE solution of the Vlasov equation

- f evolves on a timescale of order $1/\alpha$
- g evolves on a timescale of order 1

Bogoliugov "hypothesis"

We solve II supposing f fixed in time and we insert the STATIONARY solution in the r.h.s. of I

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We have reduced the problem to find the stationary solution of II equation BBGKY

$$\frac{\partial g}{\partial t} + L_f^{(1)} g + L_f^{(2)} g = C(|q_1 - q_2|) \frac{\partial f(1)}{\partial p_1} \frac{f(2)}{\partial p_2}$$

$L_f g$: linearized Vlasov operator around f acting on g

Very similar problem to solve the linear Vlasov equation

$$\frac{\partial h}{\partial t} + L_f h = 0$$

Easily doable when $f(q, p, t) = f(p, t)$

Remark: why we think that this is generalizable to fluids

$$\partial \omega + \mathbf{v} \cdot \nabla \omega = -\alpha \omega + \sqrt{\alpha} F$$

L_f --- > linear Euler operator

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KINETIC EQUATION: Non linear Fokker-Planck equation

$$\frac{1}{\alpha} \frac{\partial f(p_1, t)}{\partial t} - \frac{\partial}{\partial p_1} [p_1 f(p_1, t)] - \frac{\partial}{\partial p_1} \left[D[f](p_1) \frac{\partial f(p_1)}{\partial p_1} \right] = 0$$

 α : time-rescalingDiffusion coefficient $D[f](p_1)$

$$D[f](p) = T + 2\pi \sum_{k=1} v_k g_k \int^* dp_1 \left[\frac{1}{|\epsilon(k, kp)|^2} + \frac{1}{|\epsilon(k, kp_1)|^2} \right] \frac{1}{p_1 - p} \frac{\partial f}{\partial p} \Big|_{p_1}$$

$$T = \frac{1}{2} \sum_k g_k^2$$

$$\epsilon(k, \omega) = 1 - 2\pi i N k \varphi(k) \int_{-\infty}^{\infty} dp \frac{f'(p)}{-i\omega + ikp}$$

 $\int^* dp$: Cauchy integral v_k : Fourier components of the potential $v(q)$

Comparing the results: kinetic energy and $\langle p^4 \rangle$

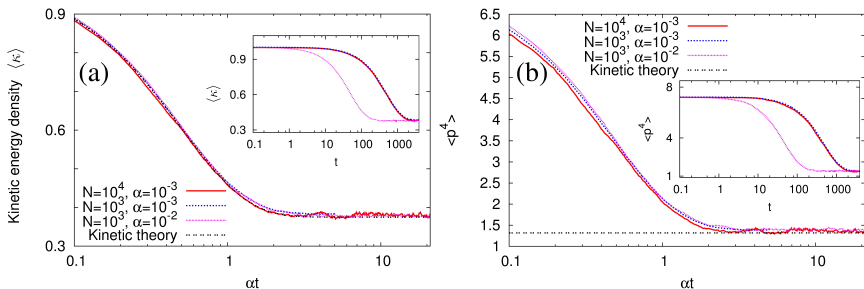
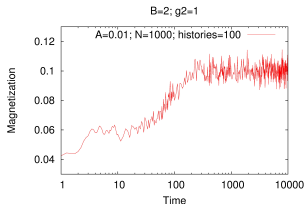
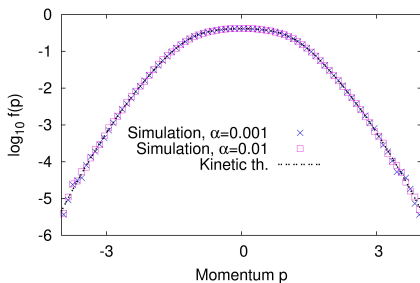
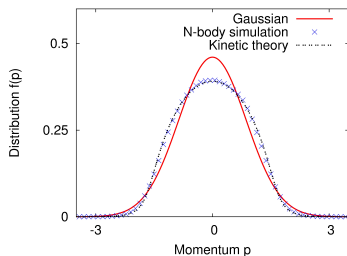


Figure: (a) Kinetic energy density $\langle \kappa \rangle$ and (b) $\langle p^4 \rangle$ as a function of αt , for the values $B_0 = 1.5$ and $g_1 = 0.75$. The data for different N and α values are obtained from numerical simulations of the stochastically forced HMF model, and involve averaging over 50 histories for $N = 10^4$ and 10^3 histories for $N = 10^3$. The data collapse implies that α is the timescale of relaxation to the stationary state. The inset shows the data without time rescaling by α .



NON-EQUILIBRIUM STATIONARY VELOCITY DISTRIBUTION

α -independent shape

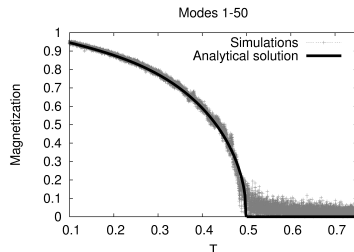


- $N = 10^4$
- kinetic temperature $T = 0.75$
- forced modes: $k = 1, 2$

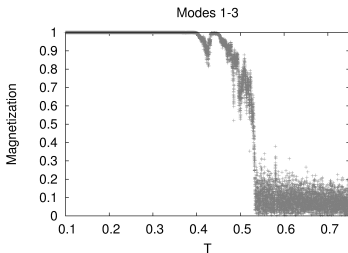
Close to detailed balance

What happens **close to the phase transition?**

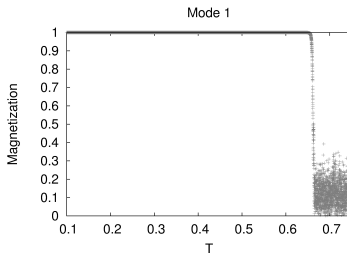
Adiabatic change of T

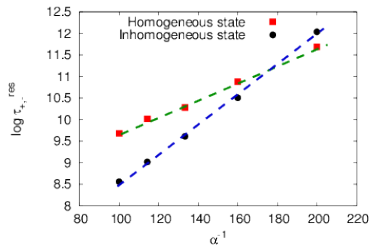
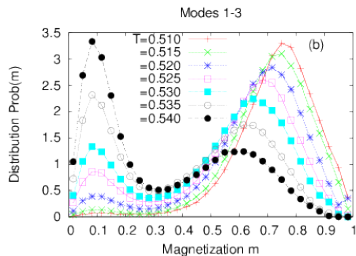
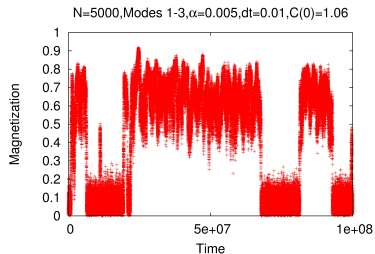


Far from detailed balance



Even further from detailed balance

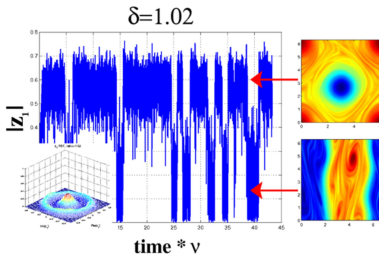




- lifetime $\sim e^{\lambda/\alpha}$
- Bistable behavior described by a Poisson process

2D Stochastic Euler

$$\partial_t \omega + \mathbf{v} \cdot \nabla \omega = -\alpha \omega + \sqrt{2\alpha} F(\mathbf{r}, t)$$



Bouchet and Simonnet, PRL, 094504 (2009)

Analogous behavior to

- Reversal of earth magnetic field
- Path of ocean currents (ex: Kuroshio current)
- Experiments on rotating fluids

Stochastically perturbed particles interacting with a long-range potential

- Kinetic theory in the weak forcing limit

Prediction of NON-EQUILIBRIUM homogeneous states

- Numerical observation of bistability

Ongoing works in kinetic theory

- 2d-turbulence: stochastic Euler equation

$$\partial_t \omega + \mathbf{v} \cdot \nabla \omega = -\alpha \omega + \sqrt{2\alpha} F(\mathbf{q}, t)$$

- with F. Bouchet and T. Tangarife

(Long term) perspectives

- kinetic theory for geophysical models?
 - 2D stochastic Navier-Stokes, Shallow-water, Quasi-geostrophic equations ...
- Theoretical understanding of the bistability

Thank you!