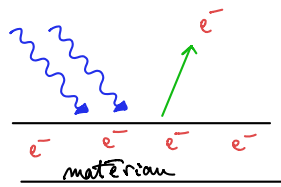
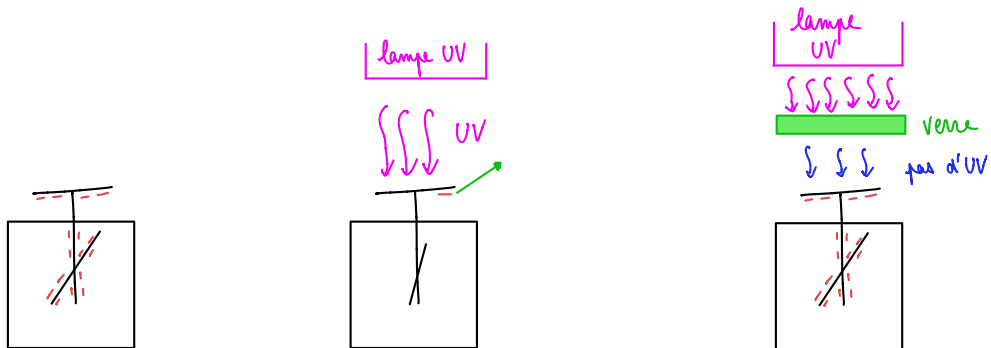
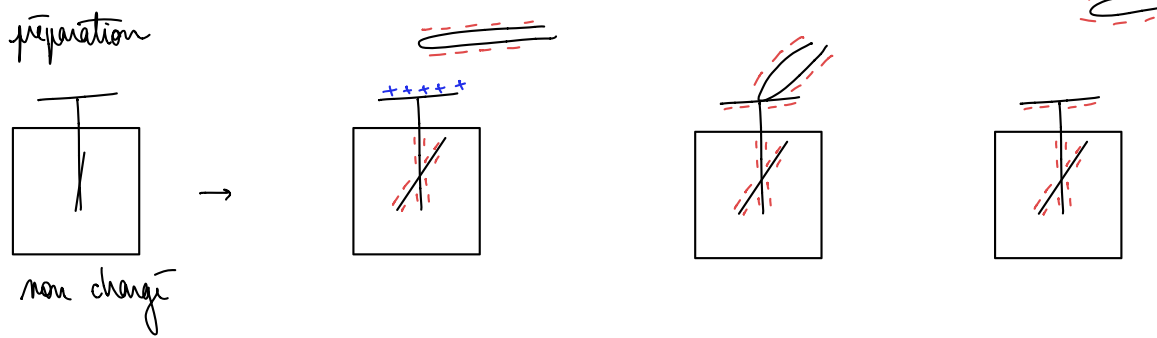
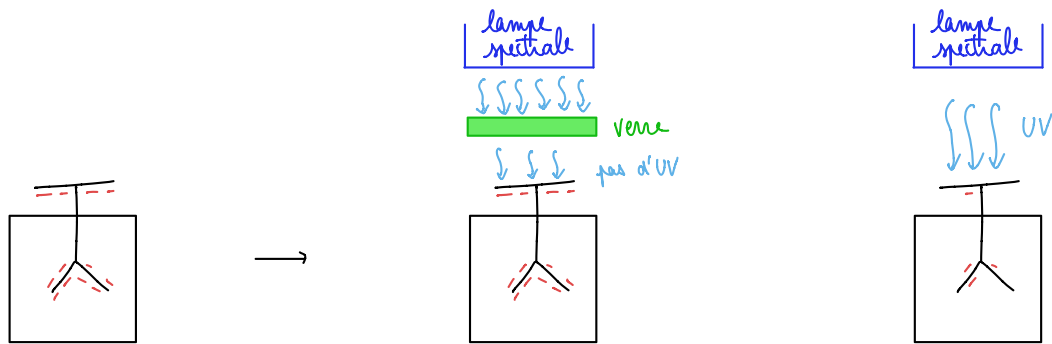
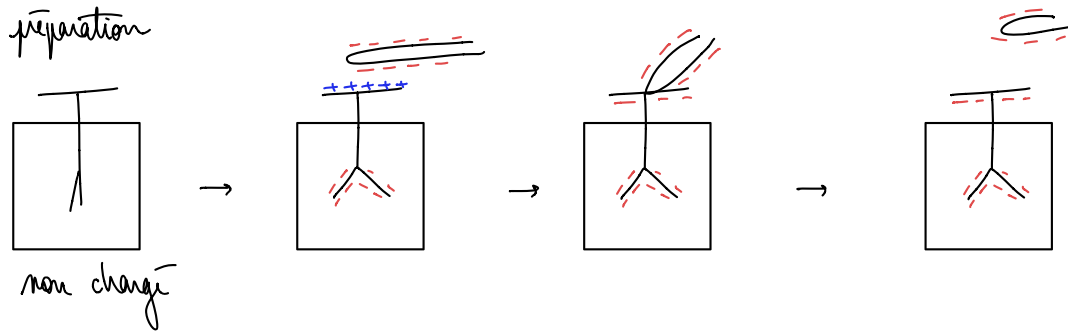
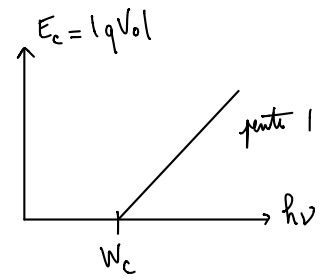
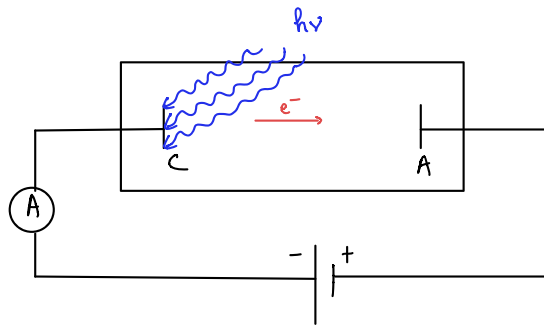


I. 1.

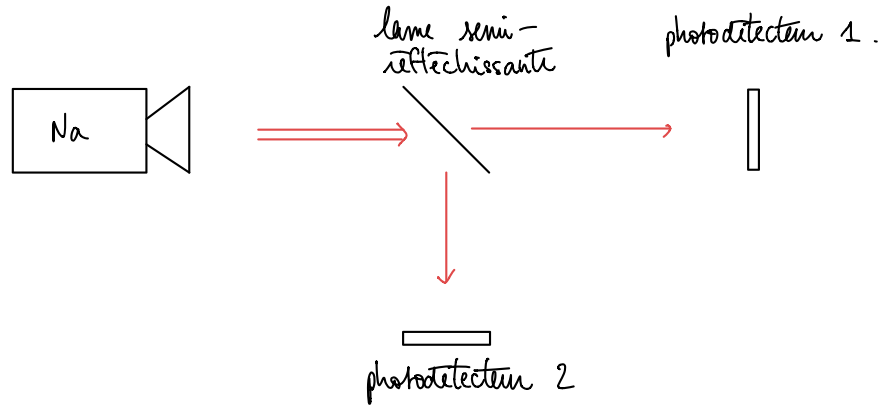


2.





3.



$$g^{(2)}(\tau) \equiv \frac{\langle \Delta I_1(t) \Delta I_2(t+\tau) \rangle}{\langle I_1(t) \rangle \langle I_2(t+\tau) \rangle}$$

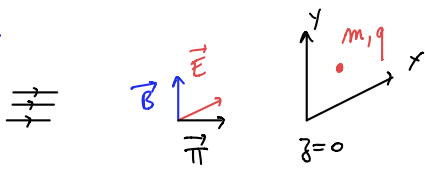
$$\text{ou } \Delta I \equiv I - \langle I \rangle$$

classique
↓

$$g^{(2)}(\tau=0) = 0.4 \rightarrow \text{anticorrelation}$$

$$(g^{(2)}(\tau) \geq 1)$$

II 1.



On considère une OPPH selon Oz arrivant sur une particule (m, q) maintenue dans le plan $z=0$

$$\vec{E}(y, t) = E_0 \cos(\omega t - k_z y) \vec{u}_y$$

$$\vec{B} = \frac{1}{c} \vec{u}_y \wedge \vec{E}$$

$$\Delta \vec{p}_{\text{Lorentz}} = \int_0^T \vec{F} \cdot d\vec{t} = \int_0^T (q \vec{v} \wedge \vec{B}(0, t)) dt = \int_0^T \frac{q}{c} \vec{v} \wedge (\vec{u}_y \wedge \vec{E}(0, t)) dt = \int_0^T \frac{q}{c} (\underbrace{\vec{v} \cdot \vec{E}(0, t)}_{\text{puissance}}) \vec{u}_y dt - \int_0^T \frac{q}{c} (\underbrace{\vec{v} \cdot \vec{u}_y}_{\text{la partie magnétique ne travaille pas}}) \vec{E}(0, t) dt$$

car particule reste à $z=0$

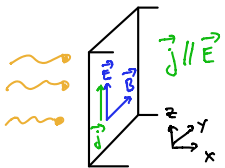
$$\Delta \vec{p} = \frac{W}{c} \vec{u}_y$$

avec les photons, $W = N h \nu$ # photons absorbés pendant 1 période

$\Delta \vec{p} = N \vec{p}_{\text{photon}}$

$\vec{p}_{\text{photon}} = \frac{h \nu}{c} \cdot \vec{u}_y$

Pression de radiation



On veut calculer P_{rad} :

$$\langle dF \rangle_t = P_{\text{rad}} dS,$$

$P_{\text{rad}} = \epsilon_0 E^2 = 2 \frac{\langle I \vec{u}_y \rangle}{c}$

$$\langle I \vec{u}_y \rangle = c \cdot \frac{\epsilon_0 E^2}{2}$$

ingrédients

$$d^2 F = d\vec{l} \wedge \vec{B}$$

$d\vec{l} = |d\vec{l}| d\vec{u}_z$ — force de Laplace

$$d\vec{l} = |d\vec{l}| dx dy, d\vec{l} = dz \vec{e}_z$$

$$\vec{j} = \frac{1}{\mu_0} \frac{\partial B_y}{\partial x} \vec{e}_z$$

— MA en négligeant $\frac{1}{c^2} \frac{\partial \vec{E}}{\partial t}$ AROS magnétique

$$B_y(x=0) = -2 \frac{E_0}{c} \cos(\omega t)$$

$$B_y(x=\infty) = 0$$

(effet de peau)

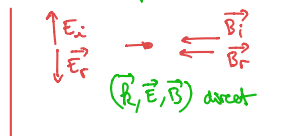
à la réflexion d'OPPH

ODG

$$P_{\text{rad}} \sim 10^{-8} \text{ Pa}$$

Puissance: laser $\sim \text{mW/mm}^2$

soleil $\sim 1 \text{ kW/m}^2 \sim \text{mW/mm}^2$ comme le laser



II)

1.

