

**1 Rayleigh wave phase velocity and error maps up to**  
**2 the fifth overtone.**

Stéphanie Durand<sup>1</sup>, Eric Debayle<sup>1</sup> and Yanick Ricard<sup>1</sup>

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Corresponding author: S. Durand, Laboratoire de Géologie de Lyon - Terre, Planète, Environnement, Université de Lyon, Bâtiment Géode, 69100 Villeurbanne, France. (stephanie.durand@ens-lyon.fr)

<sup>1</sup>Laboratoire de Géologie de Lyon - Terre, Planète, Environnement, CNRS UMR 5276, École Normale Supérieure de Lyon, Université de Lyon, Université Claude Bernard Lyon 1, 46 Allée d'Italie, 69364 Lyon Cedex 07, France

3 We present a global dataset of phase velocity maps for Rayleigh waves, with  
4 their errors. These maps are obtained from the tomographic inversion of phase  
5 velocity curves measured in the period range 40-250 s by *Debayle and Ri-*  
6 *card* [2012], completed with new measurements at longer periods, between  
7 150 and 360 s. The full dataset includes  $\sim 22,000,000$  phase velocity mea-  
8 surements combined to build 60 phase velocity maps covering the period range  
9 40-360 s for the fundamental mode and up to the fifth overtone. Each phase  
10 velocity map is provided with its *a posteriori* error, resulting in a unique dataset  
11 which can be combined with other seismic measurements (surface waves, nor-  
12 mal modes, body waves) in regional and global tomographic studies. A pre-  
13 liminary inversion of this dataset shows that it provides constraints on the  
14 shear velocity structure down to 1,000 km depth.

## 1. Introduction

15 Global seismic models suffer from a decrease of resolution in the transition zone located  
16 between  $\sim 400$  and  $\sim 1,000$  km depth. Except in very specific regions, like subduction zones  
17 [e.g. *Ritsema et al.*, 2004; *Debayle et al.*, 2005], these depths are neither well sampled by  
18 the dominant fundamental mode of surface waves nor by teleseismic body waves. It is  
19 however possible to achieve a global coverage of the transition zone by extracting more  
20 information from surface waves.

21 A first possibility is to extend the analysis to longer periods. Most datasets derived from  
22 surface waves are fundamental mode dispersion curves measured at periods shorter than  
23 200 s [*Trampert and Woodhouse*, 1995; *Ekström et al.*, 1997; *Ritzwoller and Levshin*, 1998;  
24 *Ekström*, 2011; *Visser et al.*, 2008; *Burgos et al.*, 2014] which limits the penetration depth  
25 to  $\sim 250$ -300 km. It is however possible to increase the sensitivity to the transition zone by  
26 measuring the fundamental mode at larger periods up to 360 s [*Nettles and Dziewonski*,  
27 2008].

28 A second possibility is to extract higher mode information [e.g. *Nolet*, 1975; *Cara*,  
29 1979]. Higher modes are difficult to use because they interfere in the time domain over  
30 a broad frequency range. A variety of methods for extracting higher mode dispersion  
31 curves has been developed [*Nolet*, 1975; *Cara*, 1979; *Cara and Lévêque*, 1987; *Stutzmann*  
32 *and Montagner*, 1993; *Beucler et al.*, 2003; *van Heijst and Woodhouse*, 1997; *Ekström*  
33 *et al.*, 1997; *Debayle*, 1999; *Yoshizawa and Kennett*, 2002; *Lebedev et al.*, 2005; *Visser*  
34 *et al.*, 2007]. Those using a single seismogram [*van Heijst and Woodhouse*, 1997; *Ekström*  
35 *et al.*, 1997; *Debayle*, 1999; *Yoshizawa and Kennett*, 2002; *Lebedev et al.*, 2005] achieve the

36 best coverage and have been used to built the most recent tomographic models [Ritsema  
37 *et al.*, 2011; *Debayle and Ricard*, 2012; *Schaeffer and Lebedev*, 2013].

38 In this manuscript, Rayleigh wave seismograms are analyzed with the automated wave-  
39 form inversion technique of *Debayle* [1999], modified by *Debayle and Ricard* [2012]. This  
40 waveform inversion is based on *Cara and L ev eque* [1987] and provides for a Rayleigh wave  
41 seismogram, an average 1-D velocity profile, a set of phase velocity dispersion curves and  
42 their associated *a posteriori* errors. We obtain a data base of Rayleigh wave dispersion  
43 curves for the fundamental and the five first overtones. For each mode and for each  
44 period, we have combined the measured phase velocities in a tomographic inversion to  
45 obtain phase velocity maps with their *a posteriori* errors. *A posteriori* errors provide an  
46 objective way to weight our maps when combined with other seismic datasets in future  
47 global or regional tomographic studies. A preliminary inversion of our Rayleigh wave  
48 maps alone shows that they allow to constrain mantle shear velocities down to 1,000 km  
49 depth.

## 2. Datasets

50 We use a set of dispersion curves measured by *Debayle and Ricard* [2012] from 372,629  
51 Rayleigh waveforms. Phase velocities were measured in the period range 40-250 s for the  
52 fundamental mode and up to the fifth overtone. In addition, we apply the same algorithm  
53 to 161,730 Rayleigh waveforms recorded for large events ( $6.0 < M_w < 7.2$ ) which excite  
54 surface waves at periods greater than 200 s. These waveforms are modeled in the period  
55 ranges 120-360 s, 120-240 s and 120-180 s for the fundamental, first and second higher  
56 mode, respectively.

57 The whole dataset includes  $\sim 22,000,000$  phase velocity measurements done on a total  
58 of 534,359 Rayleigh wave seismograms. A summary of the number of measurements  
59 done for each mode branch is given in Table 1. It is larger than previous Rayleigh wave  
60 datasets measured by *van Heijst and Woodhouse* [1999] or *Visser et al* [2008]. It provides  
61 new fundamental and higher mode measurements in the period range 120-360 s, which  
62 expands the fundamental mode study by *Nettles and Dziewonski* [2008]. It is comparable  
63 in size to the datasets of *Schaeffer and Lebedev* [2013] or *Ritsema et al.* [2011], but the  
64 modal and frequency contents differ. While *Schaeffer and Lebedev* [2013] analyze up to  
65 14 higher modes in the period range 10-400 s, we limit the number of modes to 6 and the  
66 period range to 40-360 s (Figure 1) which should minimize possible biases due to off great  
67 circle propagation and mode coupling effects. In addition, this dataset completes that of  
68 *Ritsema et al.* [2011] with measurements at longer periods, up to 360 s, and up to the  
69 fifth overtone.

70 Figure 1 displays the mean phase velocities and the periods and modes for which mea-  
71 surements have been made. The mean phase velocities are in general slightly larger than  
72 PREM [*Dziewonski and Anderson*, 1981] (purple dashed curves). Part of this difference  
73 may be due to the fact that a large portion of our dataset samples the thick and fast  
74 cratons of the northern hemisphere continents. The exact values used to plot Figure 1  
75 are provided in supplementary Tables 1 and 2.

76 For each individual waveform, we extract the phase velocities for the fundamental mode  
77 and up to 5 overtones and for a total of 60 periods using the iterative and non-linear  
78 inversion scheme of *Tarantola and Valette* [1982]. Our approach to fit the waveform

79 introduces a number of secondary observables [*Cara and Lévêque, 1987*]. This inversion  
80 is only weakly non-linear and we can approximate the *a posteriori* covariance matrix  
81 of our inversion at the last iteration. This processing step therefore extracts from each  
82 waveform up to 60 phase velocities with their uncertainties (the root square of the diagonal  
83 covariance matrix).

84 To decrease the number of data and improve its robustness, we cluster the seismic paths  
85 with close epicenters recorded at a given station [*Debayle and Ricard, 2012*]. We use a  
86 cluster radius of 400 km (i.e. close to the horizontal smoothing defined in section 3) and  
87 consider that the waves of a given cluster follow the same path and feel the same average  
88 dispersion between the epicenters and the station. Within each cluster, we select the  
89 best resolved phase velocity at each period and each mode, associated with the smallest *a*  
90 *posteriori* error and only when this *a posteriori* error is smaller than 80% of the *a priori*  
91 error. This path clustering allows us to restrict the whole dataset to independent paths,  
92 to discard outliers and decreases significantly the number of data and their associated  
93 uncertainties (see Table 1).

### 3. Building Phase Velocity and Error Maps

94 We combine the selected data (i.e., the phase velocities along seismics paths, measured  
95 for each mode and each period) using a continuous regionalization scheme [*Tarantola and*  
96 *Valette, 1982; Montagner, 1986; Debayle and Sambridge, 2004*]. We obtain smooth phase  
97 velocity maps by imposing correlations between neighboring points using a Gaussian *a*  
98 *priori* covariance function controlled by an horizontal correlation length, fixed to 400 km.  
99 Increasing the horizontal correlation length leads to smoother tomographic images with

100 lower amplitude anomalies, but the overall pattern of anomalies remains unchanged. An  
101 *a priori* model standard deviation controls the amplitude of the inverted model and is  
102 set to  $0.05 \text{ km s}^{-1}$ , according to the expected variation range of the shear velocity in the  
103 upper mantle [Nishimura and Forsyth, 1989; Debayle and Ricard, 2012].

104 Figure 2 presents some examples of phase velocity (middle column) and error (right  
105 column) maps. First, the phase velocity pattern is in agreement with what is expected  
106 from the sensitivity kernels (left column). For instance, for the fundamental mode at 40 s  
107 (first row) which has a maximum of sensitivity around 70 km depth, the phase velocity  
108 map principally displays the surface tectonics. If we consider a longer period, 280 s (second  
109 row), the maximum of sensitivity is now in the transition zone and the the presence of  
110 slabs become conspicuous. It can also be noted the similarity between the phase velocity  
111 maps of the fundamental mode at 40 s (first row), the third overtone at 90 s (fourth row)  
112 and the fifth overtone at 40 s (fifth row) which is related to a similar depth of maximum  
113 sensitivity. This comforts us in the quality of the measurements. Concerning the error  
114 maps (right column) they are obtained from the diagonal of the *a posteriori* covariance  
115 matrix. They are normalized to the *a priori* model standard deviation of  $0.05 \text{ km s}^{-1}$ .  
116 The shading is such that the best resolved regions are in black and the poorly resolved  
117 regions in white.

118 The whole dataset (shear velocity kernels, phase velocity maps, error maps and ray  
119 density maps) is presented in supplementary Figure 1. The error maps correlate with the  
120 ray density, the phase velocity are better constrained where the ray coverage is denser.  
121 Shear velocity sensitivity kernels show that our new dataset modeled for large events at

122 long periods (fundamental mode up to 360 s) provides additional information on the shear  
 123 velocity structure in the depth range 100-1,000 km.

#### 4. Depth inversion

124 The regionalization step leads to 60 maps of phase velocity and error (supplementary  
 125 Figure 1) for the periods and modes given in Figure 1. We extract at each geographical  
 126 point, the ensemble of phase velocities measurements and their corresponding errors that  
 127 are inverted to obtain 1D depth-dependent  $V_s$  profile. The juxtaposition of the  $V_s$  profiles  
 128 at each geographical point leads to a 3D model of shear velocities. The local phase velocity  
 129 perturbation for mode  $m$  and period  $T$ ,  $\delta c_T^m(\theta, \phi)$ , is related to the radial shear velocity  
 130 structure by :

$$131 \quad \delta c_T^m(\theta, \phi) = \int_0^a \delta V_s(\theta, \phi, r) K_T^m(r) dr, \quad (1)$$

132 where  $\delta V_s(\theta, \phi, r)$  is the local shear velocity perturbation at depth  $r$  and  $K_T^m(r) =$   
 133  $\partial c_T^m / \partial V_s(r)$  the corresponding shear wave sensitivity kernel computed in PREM (see e.g.,  
 134 Figure 2, left column). This equation implicitly assumes that the compressional velocity  
 135  $V_p$  and the density  $\rho$  are not inverted for, which is reasonable as the Rayleigh waves are  
 136 mostly affected by  $V_s$ . We also performed an inversion using scaling laws between  $V_s$  and  
 137  $V_p$  and between  $V_s$  and  $\rho$  and found a shear velocity model almost identical to the one  
 138 obtained without scaling (see supplementary Figure 3). Thus we choose to fix  $V_p$  and  
 139  $\rho$  to the values of PREM. Before inversion, we perform crustal corrections using phase  
 140 velocities predicted by the crustal part of the 3SMAC model [*Nataf and Ricard, 1996*].  
 141 The corrected phase velocity maps and the associated errors are given in supplementary  
 142 Figure 2. The uncorrected dataset is shown in supplementary Figure 1.



143 In order to get an idea of the maximal vertical resolution that we can obtain with our  
 144 dataset we compute the averaging kernels following *Backus and Gilbert* [1968]. These  
 145 averaging kernels,  $A(r, r_0)$ , relate the model which can be estimated from the inversion,  
 146  $\langle m \rangle_{r_0}$ , to the “true” model,  $m(r)$  :

$$147 \quad \langle m \rangle_{r_0} = \int_0^a A(r, r_0) m(r) dr, \quad (2)$$

148 Ideally, we would like  $A(r, r_0) = \delta(r - r_0)$  meaning that the data can resolve at each depth  
 149 an infinitely small length-scale. It is however impossible to reach such a resolution with a  
 150 finite number of data. The resolution kernels  $A(r, r_0)$  result from linear combinations of  
 151 the sensitivity kernels  $K_T^m(r)$ . They are computed by minimizing  $[A(r, r_0) - \delta(r - r_0)]^2$   
 152 [see *Backus and Gilbert*, 1968]. They are displayed in Figure 3. The shortest resolved  
 153 length-scale increases with depth, ranging from  $\sim 50$  km at 100 km depth to  $\sim 250$  km at  
 154 1500 km depth. Our dataset is barely insensitive to the structure below 1500 km depth.

155 To perform the tomographic inversion we must define the covariances matrices. The *a*  
 156 *priori* data covariance matrix is diagonal and contains the variance calculated from the  
 157 error maps shown in supplementary Figure 2. The *a priori* model covariance matrix,  $C_p$ ,  
 158 is non-diagonal: a Gaussian correlation function couples the velocity variations at radii  $r$   
 159 and  $r'$  :

$$160 \quad C_p(r, r') = \sigma^2 \exp\left(-\frac{(r - r')^2}{2L(r_M)^2}\right) \quad (3)$$

161 where  $r_M$  is  $(r + r')/2$ ,  $L$  the correlation length that controls the vertical smoothness of  
 162 the model and  $\sigma$  the *a priori* model standard deviation that controls the amplitude of  
 163 velocity variations. Based on the results of the averaging kernels we allow the correlation

length  $L$  and the *a priori* model standard deviation  $\sigma$  to vary with depth and their values  
are given in supplementary Table 3.

The shear velocity model obtained from the inversion of our phase velocity maps is displayed in Figure 4. In the upper mantle, the structures are in general agreement with other recent models derived from fundamental and higher Rayleigh modes [Debayle and Ricard, 2012; Schaeffer and Lebedev, 2013]. For example, we recover the very strong correlation with surface tectonics in the top 200 km, and the high velocity signature of ponding slabs within the transition zone. However, the addition of longer period surface waves up to 360 s provides additional sensitivity down to at least 1,000 km depths. Although these long period surface waves only constrain the broad scale structure at the top of the lower mantle, they are very complementary to other existing seismic datasets. They constrain both even and odd degrees of the shear wave mantle structure while normal mode splitting data are only sensible to the even structure. We plan to combine this surface wave dataset with normal modes and long period S-wave measurements. This will help to better constrain the mantle broad-scale structure, especially at the top of the lower mantle, where the body wave coverage remains uneven.

## 5. Concluding remarks

We present a set of fundamental and higher Rayleigh mode phase velocity maps derived from the regionalization of  $\sim 22,000,000$  phase velocity measurements. This unique dataset has been obtained from the waveform modeling of  $\sim 540,000$  waveforms (for a total of 60 periods and 6 modes), following the scheme developed by Debayle and Ricard [2012]. We provide an *a posteriori* error map associated with each phase velocity

185 map. These error maps can be used to weight our phase velocity maps in tomographic  
186 inversions combining this dataset with other seismic data in future global and regional  
187 studies. Finally, we show that an inversion of our surface wave dataset alone constrains  
188 the shear velocity structure down to at least 1,000 km depth. Both the raw measurements  
189 and the phase velocity and error maps, with and without crust corrections, are available to  
190 the community on request. This dataset could be included for the global modeling of the  
191 mantle structure, for example in the recently launched Reference Earth Model initiative.

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## References

- 197 Backus, G. and F. Gilbert, (1968), The resolving power of gross Earth data, *Geophys. J.*  
198 *R. astr. Soc.*, **16**, 169–205.
- 199 Beucler, E., E., Stutzmann, J-P., Montagner, (2003), Surface wave higher-mode phase  
200 velocity measurements using a roller-coaster-type algorithm, *Geophys. J. Int.*, *155*(1),  
201 289–307.
- 202 Burgos, G., J-P., Montagner, E., Beucler, Y., Capdeville, A., Mocquet, M., Drilleau,  
203 (2014), Oceanic lithosphere-asthenosphere boundary from surface wave dispersion data,  
204 *J. Geophys. Res.*, *119*(2), 1079–1093.

- 205 Cara, M., (1979), Lateral variations of S-velocity in the upper mantle from higher Rayleigh  
206 modes, *Geophys. J. R. Astron. Soc.*, *57*(3), 649–670.
- 207 Cara, M., J.J. Lévêque, (1987), Waveform inversion using secondary observables, *Geophys.*  
208 *Res. Lett.*, *14*, 1046–1049.
- 209 Debayle, E., (1999), SV-wave azimuthal anisotropy in the australian upper mantle: Pre-  
210 liminary results rom automated Rayleigh waveform inversion, *Geophys. J. Int.*, *137*(3),  
211 747–753.
- 212 Debayle, E., M., Sambridge, (2004), Inversion of massive surface wave data sets:  
213 Model construction and resolution assessment, *J. Geophys. Res.*, *109*, B02316,  
214 doi:10.1029/2003JB002652.
- 215 Debayle, E., B., Kennett, K., Priestley, (2004), Global azimuthal seismic anisotropy and  
216 the unique plate-motion deformation of Australia, *Nature*, *433*(7025), 509–512.
- 217 Debayle, E., and Y. Ricard, (2012), A global shear velocity model of the upper mantle from  
218 fundamental and higher Rayleigh mode measurements, *J. Geophys. Res.*, *117*(B10308),  
219 1–24.
- 220 Dziewonski, A. M., and D.I. Anderson, (1981), Preliminary reference Earth model, *Phys.*  
221 *Earth Planet. Int.*, *25*, 297–236.
- 222 Ekström, G., J., Tromp, E.W.F., Larson, (1997), Measurements and global models of  
223 surface wave propagation, *J. Geophys. Res.*, *102*(B4), 8137–8157.
- 224 Ekström, G., (2011), A global model of Love and Rayleigh surface wave dispersion and  
225 anisotropy, 25-250 s, *Geophys. J. Int.*, *187*(3), 1668–1686

- 226 Kanamori, H., Anderson, D. L., (1977), Importance of physical dispersion in surface-wave  
227 and free-oscillation problems, *Rev. Geophys.*, *15*, 105–112.
- 228 Lebedev, S., G. Nolet, T., Meier, (2005), Automated multimode inversion of surface wave  
229 and S waveforms, *Geophys. J. Int.*, *162*, 951–964.
- 230 Maupin, M., (2011), Upper-mantle structure in southern Norway from beamforming of  
231 Rayleigh wave data presenting multipathing, *Geophys. J. Int.*, *185*, 985–1002.
- 232 Montagner, J., (1986), Regional three-dimensional structures using longperiod surface  
233 waves, *Ann. Geophys.*, *4*, 283-294.
- 234 Nataf, H-C., and Y., Ricard, (1996), 3SMAC: an a priori tomographic model of the upper  
235 mantle based on geophysical modeling, *Phys. Earth Planet. Int.*, *95*, 101-122.
- 236 Nishimura, C.E., and D.W., Forsyth, (1989), The anisotropic structure of the upper  
237 mantle in the Pacific, *Geophys. J. Int.*, *96*(2), 203-229
- 238 Nettles, M., and A.M., Dziewonski, (2008), Radially anisotropic shear velocity structure  
239 of the upper mantle globally and beneath North America, *J. Geophys. Res.*, *113*(B2),  
240 1–27.
- 241 Nolet, G., (1975), Higher Rayleigh modes in western-Europe, *Geophys. Res. Lett.*, *2*(2),  
242 60–62.
- 243 Ritsema, J., H.J., van Heijst, J.H., Woodhouse, (2004), Global transition zone tomogra-  
244 phy, *J. Geophys. Res.*, *109*(B02302), 1–14.
- 245 Ritsema, J., A., Deuss, H.J., van Heijst, J.H., Woodhouse, (2011), S40RTS: a degree-40  
246 shear velocity model for the mantle from new Rayleigh wave dispersion, teleseismic  
247 traveltimes and normal-mode splitting function measurements, *Geophys. J. Int.*, *184*,

248 1223–1236.

249 Ritzwoller, M., H., and A. L. Levshin, (1998), Eurasian surface wave tomography: Group  
250 velocities, *J. Geophys. Res.*, *103*(B3), 4839–4878.

251 Shaeffer, A.J., and S., Lebedev, (2013), Global shear speed structure of the upper mantle  
252 and transition zone, *Geophys. J. Int.*, *194*(1), 417–449.

253 Stutzmann, E., and J-P., Montagner, (1993), An inverse technique for retrieving higher  
254 mode phase-velocity and mantle structure, *Geophys. J. Int.*, *113*(3), 669–683.

255 Tarantola, A., B., Valette, (1982), Generalised nonlinear inverse problems solved using  
256 the least square criterion, *Rev. Geophys.*, *20*, 219–232.

257 Trampert, J., J.H., Woodhouse, (1995), Global phase velocity maps of Love and Rayleigh  
258 waves between 40 and 150 seconds, *Geophys. J. Int.*, *122*, 675–690.

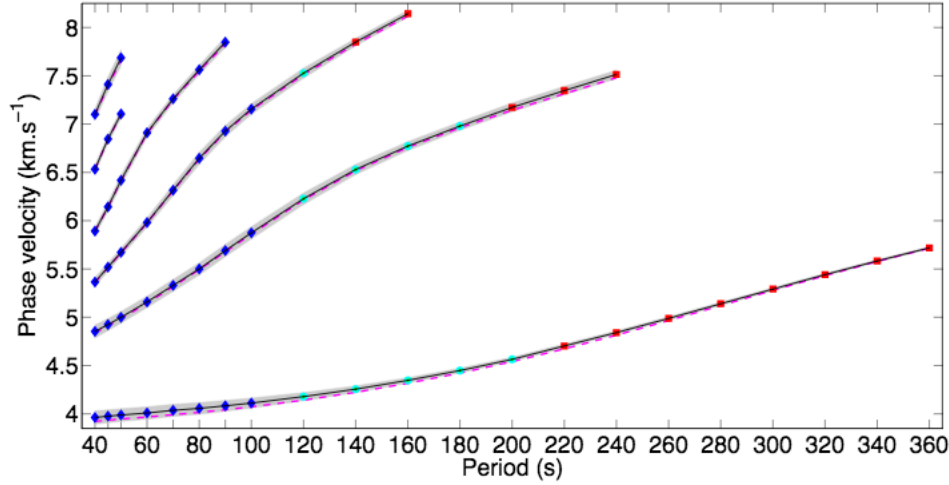
259 van Heijst, H., and J., Woodhouse, (1997), Measureing surface-wave overtone phase veloc-  
260 ities using a mode branch stripping technique, *Geophys. J. Int.*, *131*, 209–230.

261 van Heijst, H., and J., Woodhouse, (1999), Global high-resolution phase velocity distri-  
262 butions of overtone and fundamental-mode surface waves determined by mode branch  
263 stripping, *Geophys. J. Int.*, *137*, 601–620.

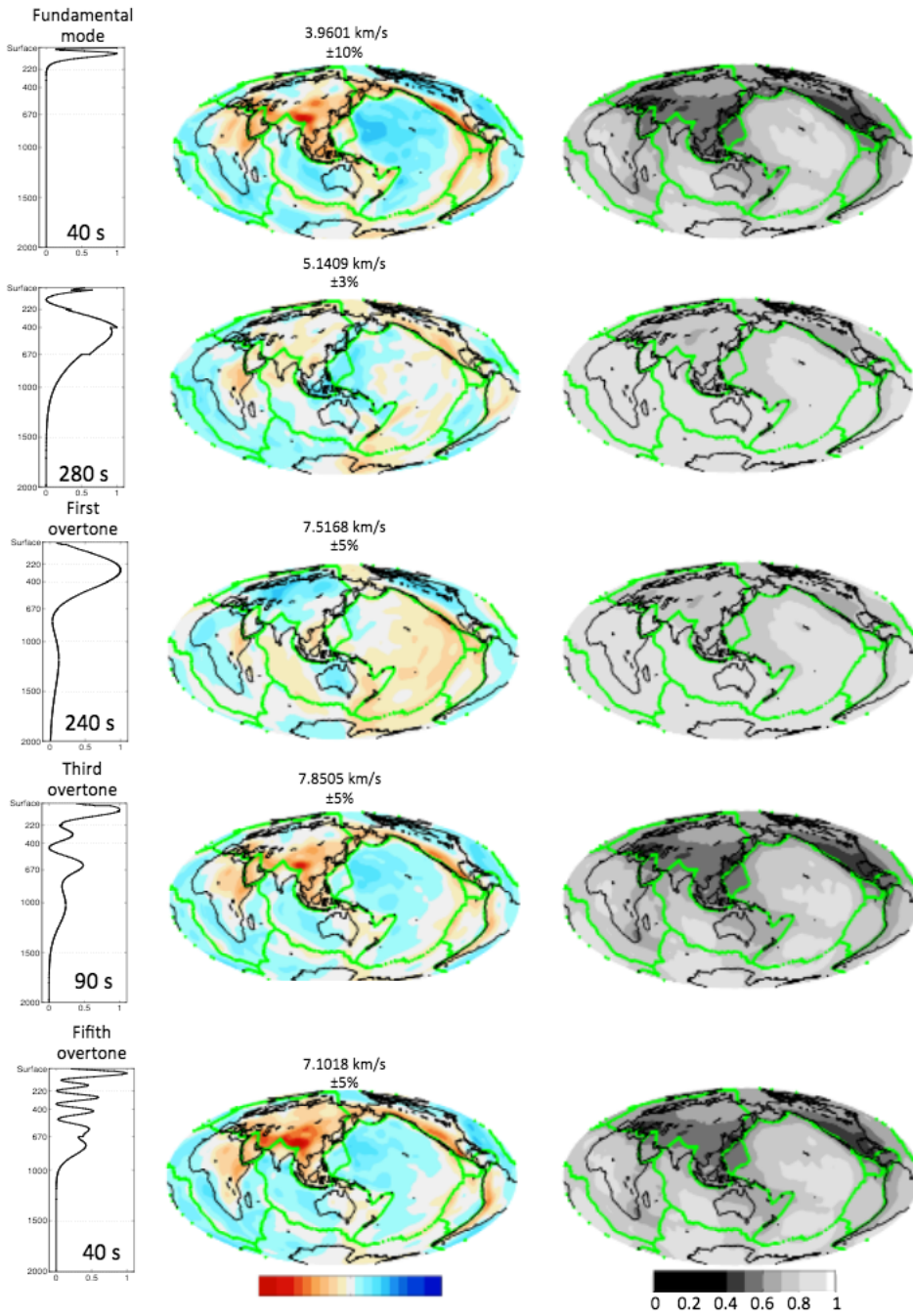
264 Visser, K., S., Lebedev, J., Trampert, B.L.N., Kennett, (2008), Global Love wave overtone  
265 measurements, *Geophys. Res. Lett.*, *34*(L03302), 11–5.

266 Visser, K., J., Trampert, B.L.N., Kennett, (2008), Global anisotropic phase velocity maps  
267 for higher ode Love and Rayleigh waves, *Geophys. J. Int.*, *172*, 1016–1032.

268 Yoshizawa, K., and B., Kennett, (2002), Determination of the influence zone for surface  
269 wave paths, *Geophys. J. Int.*, *149*, 441–454.

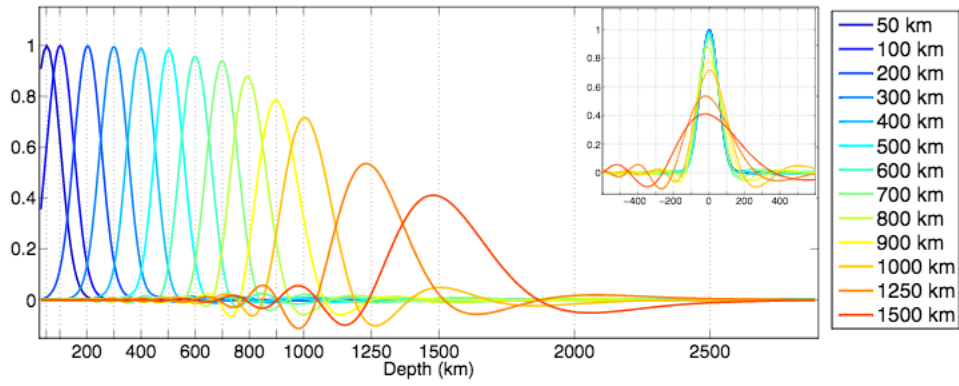


**Figure 1.** Average dispersion of our entire dataset (black curves) with one standard deviation errors (gray areas). Dispersion curves are shown for the fundamental mode and the first five overtones. Blue diamonds, light blue circles and red squares show periods and modes for which measurements are averaged over *Debayle and Ricard [2012]*'s dataset only, *Debayle and Ricard [2012]*'s dataset plus long period measurements, and only long period measurements, respectively. The dispersion of model PREM (purple dashed curves) is shown for comparison. All phase velocities are corrected from the physical dispersion due to attenuation [*Kanamori and Anderson, 1977*] using a reference period of  $T = 100$  s.

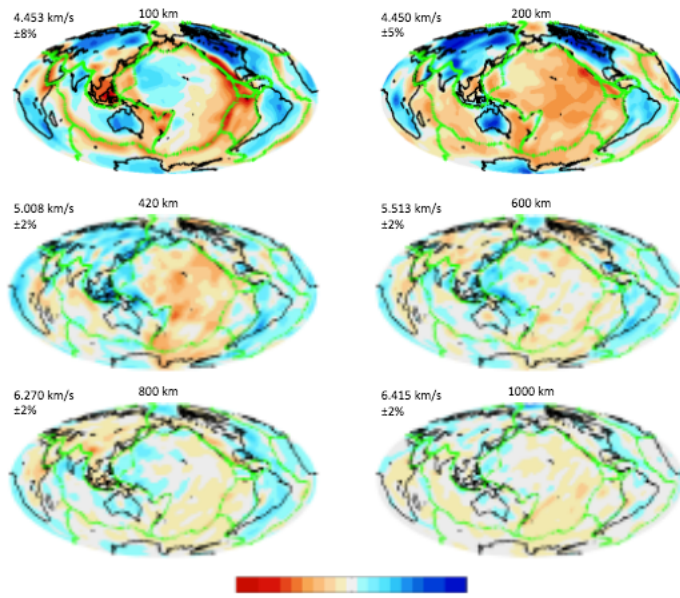


**Figure 2.** Examples of phase velocity maps (middle column) for some fundamental and higher Rayleigh modes with the corresponding error maps (right column) and  $V_S$  sensitivity kernels (left column). Velocity color scales are in percent of a reference value indicated on the top of each map. For the error map shadings, 1 corresponds to an error of 0.05 km s<sup>-1</sup>.





**Figure 3.** Averaging kernels according to *Backus and Gilbert* [1968]. Same averaging kernels aligned with respect to their maxima (Upper right corner).



**Figure 4.** 3D shear velocity model obtained from the inversion of our phase velocity maps.

Color scales are in percent of a reference value indicated on the top left corner of each map.

**Table 1.** Summary of the whole dataset indicating the number of phase velocities measured on the raw dataset and selected to built the phase velocity maps

Mode	Period range (s)	Number of measured periods	Number of raw measurements	Number of selected measurements
Fond. mode	40-360	21	7,270,127	1,559,939
1 <sup>st</sup> overtone	40-240	15	6,088,848	1,159,831
2 <sup>nd</sup> overtone	40-160	11	3,838,851	810,597
3 <sup>rd</sup> overtone	40-90	7	2,608,403	549,864
4 <sup>th</sup> overtone	40-50	3	1,117,887	214,329
5 <sup>th</sup> overtone	40-50	3	1,117,887	238,369
Total		60	22,042,003	4,532,929

